Expressions useful for discussion of Helium Description

$$
H = T_1 + V_1 + T_2 + V_2 + V_{12}
$$

 $H = T_1(\mathbf{r}_1) + V_1(\mathbf{r}_1) + T_2(\mathbf{r}_2) + V_2(\mathbf{r}_2) + V_{12}(\mathbf{r}_2, \mathbf{r}_2)$ 

$$
T_1(\mathbf{r}_1) \longrightarrow -\frac{\hbar^2}{2m_e} \nabla_{r_1}^2 \qquad T_2(\mathbf{r}_1) \longrightarrow -\frac{\hbar^2}{2m_e} \nabla_{r_2}^2
$$

$$
V_1(\mathbf{r}_1) = -\frac{Z e^2}{|\mathbf{r}_1|} \longrightarrow -\frac{Z e^2}{r_1} \qquad V_2(\mathbf{r}_2) = -\frac{Z e^2}{r_2}
$$

$$
V_{12}(\mathbf{r}_2, \mathbf{r}_2) = +\frac{e^2}{|\mathbf{r}_1 - \mathbf{r}_2|} \longrightarrow +\frac{e^2}{r_{12}}
$$

 $\Psi\left(\mathbf{r}_{1},\mathbf{r}_{2}\right)$ 

$$
H = T_1(\mathbf{r}_1) + V_1(\mathbf{r}_1) + T_2(\mathbf{r}_2) + V_2(\mathbf{r}_2) + V_{12}(\mathbf{r}_2, \mathbf{r}_2)
$$

 $[T_1(\mathbf{r}_1) + V_1(\mathbf{r}_1) + T_2(\mathbf{r}_2) + V_2(\mathbf{r}_2) + V_{12}(\mathbf{r}_2, \mathbf{r}_2)] \Psi(\mathbf{r}_1, \mathbf{r}_2) = E \Psi(\mathbf{r}_1, \mathbf{r}_2)$ 

$$
\left[ -\frac{\hbar^2}{2m_e} \nabla_{r_1}^2 - \frac{Z e^2}{r_1} - \frac{\hbar^2}{2m_e} \nabla_{r_2}^2 - \frac{Z e^2}{r_2} + \frac{e^2}{|\mathbf{r}_1 - \mathbf{r}_2|} \right] \Psi(\mathbf{r}_1, \mathbf{r}_2) = E \Psi(\mathbf{r}_1, \mathbf{r}_2)
$$

Evaluation of the repulsion term using the multipole expansion

$$
\frac{1}{|\mathbf{r}_{1}-\mathbf{r}_{2}|} = \sum_{LM} \frac{4\pi}{2L+1} \frac{r_{<}^{L}}{r_{>}^{L+1}} Y_{LM}^{\star}(\hat{r}_{1}) Y_{LM}(\hat{r}_{2})
$$
(1)

where

$$
r_{<} = r_1, \quad r_{>} = r_2 \quad \text{for} \quad |\mathbf{r}_1| < |\mathbf{r}_2|
$$
\n
$$
r_{<} = r_2, \quad r_{>} = r_1 \quad \text{for} \quad |\mathbf{r}_1| > |\mathbf{r}_2|
$$

Evaluation of the matrix element in general case

$$
\int d^3 \mathbf{r}_1 \int d^3 \mathbf{r}_2 \psi_{n_1 l_1 m_1}^{\star}(\mathbf{r}_1) \psi_{n_2 l_2 m_2}^{\star}(\mathbf{r}_2) \frac{1}{|\mathbf{r}_1 - \mathbf{r}_2|} \psi_{n_1 l_1 m_1}(\mathbf{r}_1) \psi_{n_2 l_2 m_2}(\mathbf{r}_2)
$$
 (2)

is performed separately over the radial and angular parts

$$
\int r_1^2 dr_1 \int d\hat{r}_1 \int r_2^2 dr_2 \int d\hat{r}_2 \qquad R_{n_1 l_1}^{\star}(r_1) Y_{l_1 m_1}^{\star}(\hat{r}_1) R_{n_2 l_2}^{\star}(r_2) Y_{l_2 m_2}^{\star}(\hat{r}_2)
$$

$$
\frac{1}{|\mathbf{r}_1 - \mathbf{r}_2|} \qquad R_{n_1 l_1}(r_1) Y_{l_1 m_1}(\hat{r}_1) R_{n_2 l_2}(r_2) Y_{l_2 m_2}(\hat{r}_2) \quad (3)
$$

where  $d\hat{r}_i$  means the integration over  $d\Omega_i = \sin \theta_i d\theta_i d\varphi_i$ .

The evaluation of general case - angular integrals of three  $Y_{lm}$ 's

$$
C^{L} = \int Y_{l_{i}m_{i}}^{*}(\theta, \varphi) Y_{LM}(\theta, \varphi) Y_{l_{i}m_{i}}(\theta, \varphi) d\Omega \tag{4}
$$

For the case of both s-states,  $l_i = 0$   $m_i = 0$  only  $L = 0$   $M = 0$  are nonzero; The sum reduces to one term. The angular factors give value one, since the  $(Y_{L=0M=0})^2 = (4\pi)^{-1}$  cancels the corresponding factor in the multipole expansion and due to the normalization.

Thus the repulsion matrix element with the  $e^2$  encluded

$$
\int d^{3}\mathbf{r}_{1} \int d^{3}\mathbf{r}_{2} \psi_{100}^{\star}(\mathbf{r}_{1}) \psi_{100}^{\star}(\mathbf{r}_{2}) \frac{e^{2}}{|\mathbf{r}_{1} - \mathbf{r}_{2}|} \psi_{100}(\mathbf{r}_{1}) \psi_{100}(\mathbf{r}_{2})
$$
(5)

is evaluated as the radial integral only

$$
\int r_1^2 dr_1 \int r_2^2 dr_2 R_{10}^{\star}(r_1) R_{10}^{\star}(r_2) \frac{e^2}{r_>} R_{10}(r_1) R_{10}(r_2)
$$
 (6)

# Calculating the Radial Integral

Radial Part  $R_{1,0}(r)$ :

$$
R_{1,0}(r) = 2 \cdot \left(\frac{Z}{a_0}\right)^{\frac{3}{2}} \cdot e^{-\frac{Z \cdot r}{a_0}} = R_{1,0}^*(r)
$$

Integral:

$$
\int_0^\infty \int_0^\infty r_1^2 \cdot r_2^2 \cdot R_{1,0}(r_1)^2 \cdot R_{1,0}(r_2)^2 \frac{e^2}{r_>} dr_1 dr_2
$$
  
= 
$$
\int_0^\infty \int_0^\infty 2^4 \left(\frac{Z}{a_0}\right)^6 e^{-\frac{2Z}{a_0}(r_1+r_2)} r_1^2 \cdot r_2^2 \frac{e^2}{r_>} dr_1 dr_2
$$
  
= 
$$
2^4 \left(\frac{Z}{a_0}\right)^6 \cdot e^2 \int_0^\infty \int_0^\infty e^{-\frac{2Z}{a_0}(r_1+r_2)} r_1^2 \cdot r_2^2 \frac{1}{r_>} dr_1 dr_2
$$

With substitutions  $\frac{2Z}{a_0}r_1 \to r_1$  and  $\frac{2Z}{a_0}r_2 \to r_2$ 

$$
= \frac{1}{2} \frac{Ze^2}{a_0} \underbrace{\int_0^\infty \int_0^\infty r_1^2 r_2^2 e^{-r_1} e^{-r_2} \frac{1}{r_2} dr_1 dr_2}_{int A}
$$

intA

Observe that  $\frac{e^2}{a_0}$  $\frac{e^2}{a_0} = 1a.u. = E_0$ 

To calculate the rest-integral, we split it into two integrals. For each  $r_1$  are we taking the integral over  $r_2$  and than can we take the integrale over  $r_1$ :

$$
int A = \int_0^\infty \left( \int_0^{r_1} e^{-r_1 - r_2} r_1 r_2^2 dr_2 \right) dr_1 + \int_0^\infty \left( \int_{r_1}^\infty e^{-r_1 - r_2} r_1^2 r_2 dr_2 \right) dr_1
$$
  
= 
$$
\int_0^\infty r_1 e^{-r_1} \underbrace{\int_0^{r_1} r_2^2 e^{-r_2} dr_2 dr_1}_{int B} + \int_0^\infty r_1^2 e^{-r_1} \underbrace{\int_{r_1}^\infty e^{-r_2} r_2 dr_2}_{int C} dr_1
$$

With partial integration one get:

$$
int B = 2 - e^{-r_1}(r_1^2 + 2r_1 + 2)
$$

$$
int C = e^{-r_1}(r_1 + 1)
$$

And with this you get by again merging the two split integrals:

$$
int A = \int_0^\infty 2r_1 e^{-r_1} - e^{-2r_1}(r_1^2 + 2r_1) dr_1
$$

We use

$$
\int_0^\infty x^n e^{-x} dx = n!
$$

If the exponent contains  $\alpha$ , we make substitution

$$
x = \frac{1}{\alpha}y \qquad \qquad dx = \frac{1}{\alpha} dy
$$

so that

$$
\int_0^\infty x^n dx e^{-\alpha x} = \frac{1}{\alpha^{n+1}} \int_0^\infty y^n dy e^{-y}
$$

We re-write *intA* as

$$
int A = \int_0^\infty 2r_1 e^{-r_1} dr_1 - \int_0^\infty e^{-2r_1} r_1^2 dr_1 - \int_0^\infty e^{-2r_1} 2r_1 dr_1
$$

We see that the first integral has  $n = 1$  and no constant in the exponential; thus we get 2. Second term contains  $n = 2$  and  $\alpha = 2$ . It thus gives

$$
-\frac{1}{2^3}2! = \frac{1}{4}
$$

The third term has  $n = 1$  and  $\alpha = 2$ . It gives

$$
-2\frac{1}{2^2}1! = \frac{1}{2}
$$

The final expression for

$$
A = \int_0^\infty \int_0^\infty r_1^2 r_2^2 e^{-r_1} e^{-r_2} \frac{1}{r_2} dr_1 dr_2 \tag{7}
$$

is thus

$$
A = 2 - \frac{1}{4} - \frac{1}{2} = \frac{5}{4}
$$

And with this the whole integral becomes

$$
\int d^3 \mathbf{r}_1 \int d^3 \mathbf{r}_2 \, \psi_{100}^{\star}(\mathbf{r}_1) \, \psi_{100}^{\star}(\mathbf{r}_2) \, \frac{e^2}{|\mathbf{r}_1 - \mathbf{r}_2|} \psi_{100}(\mathbf{r}_1) \, \psi_{100}(\mathbf{r}_2) = \frac{5}{8} \frac{Ze^2}{a_0} \tag{8}
$$

Acknowledgement: Alexander Sauter (PHYS261 fall 2006) has done lots of work on this part of the document

#### How to get the variational method for Helium

written by Alexander Sauter; modified by L. Kocbach ; September 2006 We start with hydrogen-like (one electron) problem

$$
H = T_1 + V_1.
$$

We remember that the kinetic energy contains only second derivatives of the wavefunction, while

$$
V_i = -\frac{Ze^2}{r_i}.
$$

We know that the ground state energy is

$$
E_{1s}(Z) = -\frac{1}{2} Z^2 \frac{e^2}{a_0}.
$$

We will need the virial theorem, in order to avoid unnecessary evaluations. It states:

$$
\left\langle T\right\rangle =-\frac{1}{2}\left\langle V\right\rangle
$$

Since

$$
\langle H \rangle = E_{1s}(Z) = -\frac{1}{2} Z^2 \frac{e^2}{a_0} = -\frac{1}{2} Z^2 E_0
$$

and

$$
\langle H\rangle=\langle T\rangle+\langle V\rangle
$$

we can see that

$$
\langle T \rangle = \frac{1}{2} Z^2 E_0
$$

and

$$
\langle V \rangle = -Z^2 \; E_0
$$

Let us now consider the variable, or unknown effective charge number  $z$ , which is contained only in the wavefunctions.

When  $z = Z$ , the kinetic energy  $\langle T \rangle$  is  $\frac{1}{2} Z^2 E_0$ . As we mentioned, the kinetic energy contains only second derivatives, no Z. That means that when z becomes different from  $Z$ , there can not be any  $Z$  in the kinetic energy  $T$ , thus

$$
\langle T(z) \rangle = \frac{1}{2} z^2 E_0
$$

On the other hand, the potential energy contains  $Z$ , as seen above. Thus

$$
\langle V(z) \rangle = - z Z E_0
$$

We look now at the total energy for two electrons including the repulsion

$$
H = T_1 + T_2 + V_1 + V_2 + V_{12}.
$$

The repulsion term  $V_{12}$  is known for the hydrogen like orbitals, or repulsion between two electrons where both are in 1s orbital. For atomic number Z we obtained

$$
V_{12} = \frac{5}{8} \frac{Ze^2}{a_0} = \frac{5}{8} Z E_0.
$$

Again, there is no  $Z$  in the repulsion energy operator, therefore

$$
V_{12}(z) = \frac{5}{8} z E_0
$$

for the orbitals with effective z.

Thus

$$
E(z) = E_0 \left(\frac{1}{2} z^2 - zZ\right) + E_0 \left(\frac{1}{2} z^2 - zZ\right) + E_0 \frac{5}{8} z.
$$
  

$$
E(z) = \left(z^2 - 2zZ + \frac{5}{8} z\right) E_0.
$$

or

The variational method says that for the ground state the energy functional

$$
E(z) = \left(z^2 - 2zZ + \frac{5}{8}z\right)E_0.
$$

must be extremal:

$$
\frac{d}{dz}E(z) = 0
$$

$$
\Leftrightarrow 2z - 2Z + \frac{5}{8} = 0
$$

$$
\Leftrightarrow z = Z - \frac{5}{16}.
$$

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#### Atomic Units

Unit of length is the Bohr radius:

$$
a_0 = \frac{\hbar^2}{m_e e^2} \left( \frac{}{}\,\, = 4\pi\epsilon_0 \frac{\hbar^2}{m_e e^2} \, \, \right)
$$

The first is in atomic units, second in SI-units. This quantity can be remembered by recalling the virial theorem, i.e. that in absolute value, half of the potential energy is equal to the kinetic energy. This gives us

$$
\frac{1}{2}\frac{e^2}{a_0} = \frac{\hbar^2}{2m_e a_0^2}
$$

and if we accept this relation, we have the above value of  $a_0$ .

The so called fine structure constant

$$
\alpha = \frac{e^2}{\hbar c}
$$

expresses in general the weakness of electromagnetic interaction.

### Some Constants and Quantities



$$
\mu_B = 0.579 \ 10^{-4} \ \text{eV} \ (\text{ Tesla})^{-1}
$$
 Bohr magneton

#### Plank's formula

$$
\rho(\omega_{ba}) = \frac{\hbar \omega_{ba}^3}{\pi^2 c^3} \frac{1}{e^{\hbar \omega/kT} - 1}
$$

## Useful formulae and informations

$$
P_0(\cos \theta) = 1 \qquad \qquad P_1(\cos \theta) = \cos \theta
$$